# Flow field stratigraphy in north Mahuea Tholus quadrangle (V-49), Venus Rogers, KayLeigh A., Covley, Margaret T., Lang, Nicholas P. and Thomson, Bradley James

## 1) Intro

The northern third of the Mahuea Tholus quadrangle of Venus (V49; 25-50°S, 150-180°E) hosts Diana-Dali Chasma – a significant ~NE-SW trending zone of extension. Emanating from this zone are multiple lava flows that appear to be sourced from both coronae and individual fractures. Flow materials have traveled both north and south of this zone, creating an extensive flow field that encompasses an area of 1.2 x 10<sup>6</sup> km<sup>2</sup> across the northern and central portions of V-49. As part of an effort to constrain the geologic history of the V-49 quadrangle, we have been unraveling the stratigraphy of the flow field associated with this rift zone. In the surrounding quadrangles, the Diana-Dali Chasma has been divided by major coronaassociated flow fields that continue into V-49. These flows are identified as the central points of volcanism within the rift zone and were used to help in determining stratigraphy. Specifically, we have utilized NASA Magellan SAR imagery (~100 m/pxl) and altimetry data (~5 km/pxl) to geologically map individual flow units and their associated features.

# 2) Methodology

- NASA Magellan SAR Imagery (~100 m/pxl)
- Altimetry Data (~5 km/pxl)
- Hansen (2000) Methodology
- ArcMap





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Description of Map Units							
ba	<b>Basal Material, unit a</b> – Bright to Moder- ately Dark on SAR images; hosts abundant short NE to SW trending wrinkle ridges	сМа	<b>Mayauel Coronae flow</b> - Moderately bright on SAR images; flow indicators mapped on the basis of lobate and digitate flow bound- aries provide indications of local source regions; radial or concentric fractures locally deform the flows; low viscosity flows emplaced following preexisting topography	<mark>cAn</mark>	<b>Annapurna Coronae Flow</b> – Moderate- ly bright on SAR images; flow indicators mapped on the basis of digitate flow boundar- ies provide indications of local source regions; also, provide evidence that in general these flows emanate from fractures that are radial or concentric	cAg	P e c fl F
bb1	<b>Basal Material, unit 1b</b> – Bright on SAR images; hosts abundant fractures extensively cut with two dominate trends (NW and SE), and broad wrinkle ridges of two dominate trends (NE and SW); No lava flows within the false tesserae area	cC	<b>Covered Coronae</b> - Moderately bright on SAR; low viscosity flows; NW to SE trending frac- tures show evidence of a circular fracture zone, evidence of a covered coronae is visible	cFl	<b>Flidais Coronae Flow</b> – Moderately dark on SAR images; in general these flows emanate from fractures flow to SW and South from radial fractures in the Northern quadran- gle; radial or concentric fractures locally deform the flows; low viscosity flows	aMtf	N S L S
bb2	<b>Basal Material, unit 2b</b> – Bright on SAR images; hosts abundant short NE to SW trending fractures	fDd	<b>Diana-Dali Chasma flow</b> - Moderately bright on SAR images; flow originates from within the chasma then follows topo trends into the	fzDd	<b>Diana-Dali Chasma, fracture zone</b> - Bright on SAR images; Conglomeration of lava flows erupted from fractures associated with	cUe	() () ()

shield field low plain



Agraulos Coronae flow – Bright to moderately dark on SAR images; flow indicators provide evidence that in general these flows emanate from fractures that are either radial o concentric and locally deforming; low viscosit lows emplaced following preexisting topogra

Mahuea Tholus flow, unit a – Bright on SAR Images; NE – trending wrinkle ridges are underlying, relatively smooth layered planar surface; highly viscous flow

Crater Material, undifferentiated ejecta – Very bright on SAR images; includes floor, central peak, wall, rim, and ejecta material; Deposits and structures created by meteorite impact

multiple coronae in the area; general fracture

direction is NE to SW

- Atahensik Coronae flow Moderately bright on SAR images; in general these flows emanate from fractures that are either radial or concentric; radial or concentric fractures locally deform the flows; low viscosity flows emplaced following preexisting topography
- Mahuea Tholus flow, unit b Moderate-Iy bright on SAR Images; NE – trending wrinkle ridges amd NW fractures are underlying, relatively smooth planar surface; highly viscous flow
- Crater Material, undifferentiated **flow** – Bright to Dark on SAR images; includes exterior flood lava materials; texture granular to smooth. Deposits and structures created by meteorite impact

**Colijnsplaat Coronae Flow** – Moderately bright on SAR images; flow boundaries provide indications of local source regions; in general these flows emanate from fractures that are either radial or concentric

**Shield flow material** – Bright to moderately bright on SAR images; mottled texture; generally circular flow patterns







### References

V.L. Hansen, H.R. DeShon. Geologic Map of the Diana-Dali Chasma Quadrangle (V-37), Venus N.P. Lang, V.L. Hansen. Geologic Map of the greenway Quadrangle (V-24), Venus **V.L. Hansen.** *Geologic mapping of tectonic planets (2000)* 

### Acknowedgements



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