

1. Motivation

sub-solidus conditions (Challener and Glazner, 2017; Ackerson et al. 2018) of granitoid rocks in the Sierra Nevada batholith (SNB). Such discoveries beg deeper investigation of the transition from magmatic to sub-solidus conditions in arc batholiths and hold potential for understanding their changing thermal state during the waxing and waning of magmatism. Understanding of this sub-solidus realm also bears on the fluid mediated formation of ore deposits and modulation of volatile fluxes (e.g., CO2). In this work, we have examined Pb-U systematics and trace ry of both primary, euhedral titanite and co-existing, texturally late, from ~30 rocks in the western, central, and eastern Sierra Nevada, ought to 1) evaluate relative Pb-U ages of titanite and zircon; 2) to probe the chemistry of secondary titanite as a recorder of the sub-solidus history o individual rocks as well as trends across the batholit





Figures 1&2: Simplified geologic map showing the regional extent of the SNB and relative ages of plutonic and wallrock units; study area shown with dashed red line (left). Map of study area intrusive suites and their respective age ranges, based on Lackey et al. (2008) and Coleman et al. (2004).

2. Sierra Nevada Titanite



Figure 3: Textures of primary and secondary titanite. Primary titanite in top row shows euhedral look and is associated with feldspar, but often occurs in association with hornblende and/or biotite. Backscattered electron (BSE) image showing in second row, showing oscillatory zoning in primary and sector zoning in primary titanite crystals. Middle-right sample from 1S91 showing color difference between primary prown-orange titanite and pale, secondary titanite from same rock. In third row, petrographic details of secondary titanite growth which typically is finer grained and forms coronas around oxides and where biotite is altered to chlorite.

Experiments: Primary and secondary titanite was hand picked from concentrates Pb-U ages were measured at UC Santa Barbara after Kylander-Clark et al. (2013) and at Pomona College's Oxtoby Isotope Lab via laser ablation on an Agilent 8900 triple quadrupole mass spectrometer in He cell gas configuration. Titanite BLR and Y17 were employed as primary and secondary standards at UCSB. The Pomona lab employed MKED-1 and BLR-1 as primary standards for U-Pb and trace element ncentrations, respectively, Previous itanium EMPA results of typical Sierra titanite were used as internal elementa andard values. NIST 612 was also analyzed as a validation standard for trace element determinations. Data was reduced reduction software.



TITANITE U-PB AGE AND TRACE ELEMENT CHEMISTRY: TRACING COOLING AND ALTERATION IN THE SIERRA NEVADA BATHOLITH HRUSKA, Grace, LACKEY, Jade Star and MCCARTY, Kyle R.,

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Figure 5: Correspondence of age for zircon and titanite shows variable offset to younger age of titanite compared to zircon, although often within error of the 1:1 dashed line which indicates analogous age. Some cases of titanite are up to 15 million years younger than zircon. Samples of Chen and Moore (1982) shown or reference

Pb-U Titanite (Ma)







Figure 9: Titanite Tera-Wasserburg Concordia age diagram, REE plot, and histogram illustrating two different titanite populations from Sample 17GH3. Majority of the titanite within the samples analyzed are igneous (magmatic) and have trace element pattern consistent with magmatic growth. Secondary titanite has lower total REEs and posi-Eu anomalies suggesting growth in equilibrium with deuteric fluids.

Figure 10: Variations of δ^{18} O in primary and secondary titanite with respect to δ^{18} O of zircon. Inset diagrams depict scenarios of expected δ^{18} O variation: (A) differing magmatic δ^{18} O, (B) variable temperature of growth (or diffusive exchange), and (C) shifts expected from meteoric water exchange. Note the very low δ^{18} O values (< 2‰) of secondary titanite for two samples. Overall, a dearth of unusually low δ^{18} O values suggests that incursion of exogenous or meteoric water was limited, but deuteric fluid alteration, commonly indicated by retrogression of mafic minerals and feldspars, was likely widespread. Data compiled from Lackey et al. (2008) and Lackey (2005). Isotherms after King et al. (2001).





7. Key Points

- Titanite ages often post-date U-Pb zircon, confirming patterns originally identified by Chen and Moore (1982); higher precision geoand thermochronology will be needed to establish how cooling regime affects apparent age offsets.
- Secondary titanite shows distinct budget-limited depletion of REEs and greater variety in overall chemistry reflecting a mix of closed and open-system processes over a range of temperatures.
- Small, older plutons adjacent to large, late Cretaceous intrusive suites show lag of titanite U-Pb ages by up to 10 m.y., potentially a record of protracted growth, or re-crystallization (Schwartz et al. 2016), or potential thermal resetting.
- On a cautionary note, some late-stage titanite post-dates igneous crystallization ages by 10s of millions of years which warrants caution when interpreting detrital titanite U-Pb ages, in absence of REE patterns.
- Ongoing work is assessing intracrystalline patterns of Pb/U age and trace element concentration to establish if diffusional resetting or multiple growth episodes (e.g. Schwartz et al. 2016) explain marked age discrepancies in some titanite with igneous REE patterns.
- Overall, we note that Sierra Crest plutons have closer agreement of U-Pb age of zircon and titanite, suggesting that younger plutons saw a more rapid cooling, thus greater match of zircon and titanite U-Pb age. Western plutons, which may have seen titanite grow in response to changes of redox state in the sub-solidus, do not present a clear signal between diffusional Pb loss. We infer that growth of substantially younger (up to 10 million years) titanite was initiated as younger magmas caused re-equilibration of Fe-Ti oxide minerals with the younger deuteric fluids.

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